

Mössbauer neutrinos

Joachim Kopp

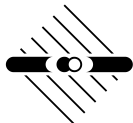
Max-Planck-Institut für Kernphysik, Heidelberg

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JHEP **0805** (2008) 005 (arXiv:0802.2513),
J. Phys. **G 36** (2009) 078001 (arXiv:0803.1424)
JHEP **0906** (2009) 049 (arXiv:0904.4346)



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Outline

- 1 The Mössbauer neutrino experiment
- 2 Oscillations of Mössbauer neutrinos: Qualitative arguments
- 3 Mössbauer neutrinos in QFT
- 4 Conclusions

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Classical Mössbauer effect

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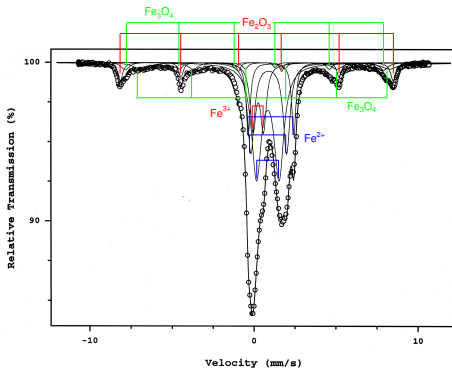
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Mössbauer neutrinos

A similar effect should exist for neutrino emission/absorption in bound state β decay and induced electron capture processes.

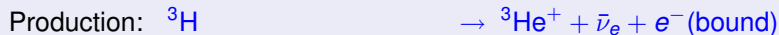
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Proposed experiment:



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Proposed experiment:

Production: ${}^3\text{H} \rightarrow {}^3\text{He}^+ + \bar{\nu}_e + e^- (\text{bound})$

Detection: ${}^3\text{He}^+ + e^- (\text{bound}) + \bar{\nu}_e \rightarrow {}^3\text{H}$

${}^3\text{H}$ and ${}^3\text{He}$ are embedded in metal crystals (metal hydrides).

Physics goals:

- Neutrino oscillations on a laboratory scale: $E = 18.6 \text{ keV}$, $L_{\text{atm}}^{\text{osc}} \sim 20 \text{ m}$.
- Gravitational interactions of neutrinos
- Study of solid state effects with unprecedented precision

Mössbauer neutrinos (2)

Mössbauer neutrinos have very special properties:

- Neutrino receives *full* decay energy: $Q = 18.6 \text{ keV}$
- Natural line width: $\gamma \sim 1.17 \times 10^{-24} \text{ eV}$
- Actual line width: $\gamma \gtrsim 10^{-11} \text{ eV}$
 - ▶ Inhomogeneous broadening (Impurities, lattice defects)
 - ▶ Homogeneous broadening (Spin interactions)

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Experimental challenges:

- Is the Lamb-Mössbauer factor (fraction of recoil-free emissions/absorptions) large enough?
- Can a linewidth $\gamma \gtrsim 10^{-11}$ eV be achieved?
- Can the resonance condition be fulfilled?

Mössbauer neutrinos (3)

Recent controversy:

- Does the small energy uncertainty prohibit oscillations of Mössbauer neutrinos?
- Do oscillating neutrinos need to have equal energies resp. equal momenta?

S. M. Bilenky, F. v. Feilitzsch, W. Potzel, J. Phys. **G34** (2007) 987, hep-ph/0611285

- Does the time-energy uncertainty relation prevent oscillations?

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⇒ Careful treatment with as few assumptions as possible is needed

⇒ Answer to the above questions will be No.

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Textbook derivation of the oscillation formula

Diagonalization of the mass terms of the charged leptons and neutrinos gives

$$\mathcal{L} \supset -\frac{g}{\sqrt{2}} (\bar{e}_{\alpha L} \gamma^\mu U_{\alpha j} \nu_{jL}) W_\mu^- + \text{diag. mass terms} + h.c.$$

(flavour eigenstates: $\alpha = e, \mu, \tau$, mass eigenstates: $j = 1, 2, 3$)

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Assume, at time $t = 0$ and location $\vec{x} = 0$, a flavour eigenstate

$$|\nu(0, 0)\rangle = |\nu_\alpha\rangle = \sum_i U_{\alpha j}^* |\nu_j\rangle$$

is produced. At time t and position \vec{x} , it has evolved into

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Oscillation probability:

$$P(\nu_\alpha \rightarrow \nu_\beta) = \left| \langle \nu_\beta | \nu(t, \vec{x}) \rangle \right|^2 = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* e^{-i(E_j - E_k)t + i(\vec{p}_j - \vec{p}_k) \cdot \vec{x}}$$

Equal energies or equal momenta?

Typical *assumptions* in the “textbook derivation” of the oscillation formula:

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These are *assumptions* or *approximations*, not fundamental principles!

Problems with the textbook derivation

- In general, neither the equal energy assumption nor the equal momentum assumption is physically justified because both violate energy-momentum conservation in the production and detection processes.

R. G. Winter, Lett. Nuovo Cim. **30** (1981) 101

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Energy-momentum conservation for emission of mass eigenstate $|\nu_i\rangle$:

$$E_i^2 = \frac{m_\pi^2}{4} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right)^2 + \frac{m_i^2}{2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right) + \frac{m_i^4}{4m_\pi^2}$$

$$p_i^2 = \frac{m_\pi^2}{4} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right)^2 - \frac{m_i^2}{2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right) + \frac{m_i^4}{4m_\pi^2}$$

For massless neutrinos: $E_i = p_i = E \equiv \frac{m_\pi}{2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right) \simeq 30 \text{ MeV}$.

To first order in m_i^2 :

$$E_i \simeq E + \xi \frac{m_i^2}{2E}, \quad p_i \simeq E - (1 - \xi) \frac{m_i^2}{2E}, \quad \xi \approx \frac{1}{2} \left(1 - \frac{m_\mu^2}{m_\pi^2}\right) \approx 0.2$$

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 - ▶ Requires neither equal E nor equal p
 - ▶ Takes into account finite resolutions of the source and the detector

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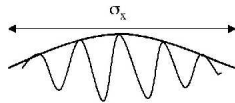
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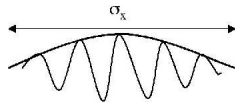
Beuthe, Giunti, Grimus, Kiers, Kim, Lee, Mohanty, Nussinov, Stockinger, Weiss, ...

Conditions for oscillations in a wave packet approach



Conditions for oscillations in a wave packet approach

- Coherence in production and detection processes

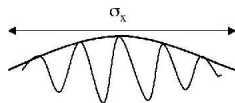


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Neutrino oscillations are caused by the superposition of different mass eigenstates.

- ⇒ If an experiment can distinguish different mass eigenstates, oscillations will vanish.

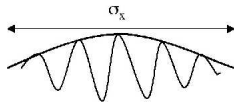


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Requirement for mass resolution σ_m :

$$\sigma_m^2 = \sqrt{(2E\sigma_E)^2 + (2p\sigma_p)^2} > \Delta m^2$$

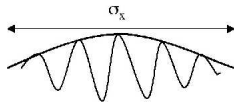
B. Kayser, Phys. Rev. **D24** (1981) 110

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This is easily fulfilled for Mössbauer neutrinos, since

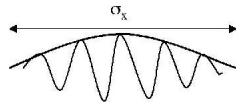
$$\sigma_E \sim 10^{-11} \text{ eV}$$

$$\sigma_p = 1/2\sigma_x \sim 1/\text{interatomic distance} \sim 10 \text{ keV}$$

$$E = p = 18.6 \text{ keV}$$

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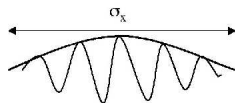
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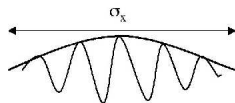
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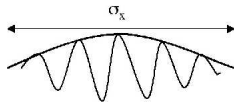
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$$L^{\text{osc}} \ll L^{\text{coh}}.$$

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It can be shown that, for Mössbauer neutrinos, σ_p is small enough, so that

$$L^{\text{osc}} \ll L^{\text{coh}}.$$

⇒ Standard oscillation formula is approximately recovered:

$$P_{ee} = \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp \left[-2\pi i \frac{L}{L_{jk}^{\text{osc}}} \right]$$

$$L_{jk}^{\text{osc}} = \frac{4\pi E}{\Delta m_{jk}^2}$$

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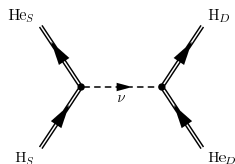
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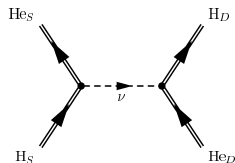
Idea: Treat neutrino as an internal line in a tree level Feynman diagram:



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Idea: Treat neutrino as an internal line in a tree level Feynman diagram:



External particles reside in harmonic oscillator potentials.
E.g. for ${}^3\text{H}$ atoms in the source:

$$\psi_{H,S}(\vec{x}, t) = \left[\frac{m_H \omega_{H,S}}{\pi} \right]^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_H \omega_{H,S} |\vec{x} - \vec{x}_S|^2 \right] \cdot e^{-iE_{H,S}t}$$

Oscillation amplitude

$$\begin{aligned}
 i\mathcal{A} = & \int d^3x_1 dt_1 \int d^3x_2 dt_2 \left(\frac{m_{H\omega_{H,S}}}{\pi} \right)^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_{H\omega_{H,S}} |\vec{x}_1 - \vec{x}_S|^2 \right] e^{-iE_{H,S}t_1} \\
 & \cdot \left(\frac{m_{He\omega_{He,S}}}{\pi} \right)^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_{He\omega_{He,S}} |\vec{x}_1 - \vec{x}_S|^2 \right] e^{+iE_{He,S}t_1} \\
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 & \cdot \left(\frac{m_{H\omega_{H,D}}}{\pi} \right)^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_{H\omega_{H,D}} |\vec{x}_2 - \vec{x}_D|^2 \right] e^{+iE_{H,D}t_2} \\
 & \cdot \sum_j \mathcal{M}^\mu \mathcal{M}^{\nu*} |U_{ej}|^2 \int \frac{d^4p}{(2\pi)^4} e^{-ip_0(t_2-t_1) + i\vec{p}(\vec{x}_2-\vec{x}_1)} \\
 & \cdot \bar{u}_{e,S} \gamma_\mu (1 - \gamma^5) \frac{i(\not{p} + m_j)}{p_0^2 - \vec{p}^2 - m_j^2 + i\epsilon} (1 + \gamma^5) \gamma_\nu u_{e,D}.
 \end{aligned}$$

Oscillation amplitude

$$\begin{aligned}
 i\mathcal{A} = & \int d^3x_1 dt_1 \int d^3x_2 dt_2 \left(\frac{m_{H\omega_{H,S}}}{\pi} \right)^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_{H\omega_{H,S}} |\vec{x}_1 - \vec{x}_S|^2 \right] e^{-iE_{H,S}t_1} \\
 & \cdot \left(\frac{m_{He\omega_{He,S}}}{\pi} \right)^{\frac{3}{4}} \exp \left[-\frac{1}{2} m_{He\omega_{He,S}} |\vec{x}_1 - \vec{x}_S|^2 \right] e^{+iE_{He,S}t_1} \\
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 \end{aligned}$$

Evaluation:

- $dt_1 dt_2$ -integrals \rightarrow energy-conserving δ functions $\rightarrow p_0$ -integral trivial
- $d^3x_1 d^3x_2$ -integrals are Gaussian
- d^3p -integral: Use Grimus-Stockinger theorem

The Grimus-Stockinger theorem

Let $\psi(\vec{p})$ be a three times continuously differentiable function on \mathbb{R}^3 , such that ψ itself and all its first and second derivatives decrease at least like $1/|\vec{p}|^2$ for $|\vec{p}| \rightarrow \infty$. Then, for any real number $A > 0$,

$$\int d^3p \frac{\psi(\vec{p}) e^{i\vec{p}\vec{L}}}{A - \vec{p}^2 + i\epsilon} \xrightarrow{|\vec{L}| \rightarrow \infty} -\frac{2\pi^2}{L} \psi(\sqrt{A}\frac{\vec{L}}{L}) e^{i\sqrt{A}L} + \mathcal{O}(L^{-\frac{3}{2}}).$$

⇒ Quantification of requirement of on-shellness for large $L = |\vec{L}|$.

W. Grimus, P. Stockinger, Phys. Rev. **D54** (1996) 3414, hep-ph/9603430

From the amplitude to the transition rate

Amplitude:

$$i\mathcal{A} = \frac{-i}{2L} \mathcal{N} \delta(E_S - E_D) \exp\left[-\frac{E_S^2 - m_j^2}{2\sigma_p^2}\right] \sum_j \mathcal{M}^\mu \mathcal{M}^{\nu*} |U_{ej}|^2 e^{i\sqrt{E_S^2 - m_j^2}L} \\ \cdot \bar{u}_{e,S} \gamma_\mu \frac{1-\gamma^5}{2} (\not{p}_j + m_j) \frac{1+\gamma^5}{2} \gamma_\nu u_{e,D},$$
$$\sigma_p^{-2} = (m_H \omega_{H,S} + m_{He} \omega_{He,S})^{-1} + (m_H \omega_{H,D} + m_{He} \omega_{He,D})^{-1}$$

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Transition rate: Integrate $|\mathcal{A}|^2$ over densities of initial and final states

$$\Gamma \propto \int_0^\infty dE_{H,S} dE_{He,S} dE_{He,D} dE_{H,D}$$

$$\cdot \delta(E_S - E_D) \rho_{H,S}(E_{H,S}) \rho_{He,D}(E_{He,D}) \rho_{He,S}(E_{He,S}) \rho_{H,D}(E_{H,D})$$

$$\cdot \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \underbrace{\exp\left[-\frac{2E_S^2 - m_j^2 - m_k^2}{2\sigma_p^2}\right]}_{\text{Analogue of Lamb-Mössbauer factor (Recoil-free fraction)}} \underbrace{e^{i(\sqrt{E_S^2 - m_j^2} - \sqrt{E_S^2 - m_k^2})L}}_{\text{Oscillation phase}}$$

The Lamb-Mössbauer factor

The **Lamb-Mössbauer factor** is the relative probability of recoil-free emission and absorption, compared to the total emission and absorption probability.

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Convenient reformulation:

$$\exp \left[- \frac{2E_S^2 - m_j^2 - m_k^2}{2\sigma_p^2} \right] = \exp \left[- \frac{(p_{jk}^{\min})^2}{\sigma_p^2} \right] \exp \left[- \frac{|\Delta m_{jk}^2|}{2\sigma_p^2} \right]$$

where $(p_{jk}^{\min})^2 = E_S^2 - \max(m_j^2, m_k^2)$.

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where $(p_{jk}^{\min})^2 = E_S^2 - \max(m_j^2, m_k^2)$.

⇒ **Localization condition**

$$4\pi\sigma_x E / \sigma_p \lesssim L_{jk}^{\text{osc}},$$

(with $\sigma_x = 1/2\sigma_p$) is satisfied if $L_{jk}^{\text{osc}} \gtrsim 2\pi\sigma_x$, which is easily fulfilled in realistic situations.

Line broadening

Energy levels of ^3H and ^3He in the source and detector are smeared e.g. due to spin-spin interactions, crystal impurities, lattice defects, etc.

R. S. Raghavan, hep-ph/0601079

W. Potzel, Phys. Scripta **T127** (2006) 85

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$$\rho_{A,B}(E_{A,B}) = \frac{\gamma_{A,B}/2\pi}{(E_{A,B} - E_{A,B,0})^2 + \gamma_{A,B}^2/4}$$

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Result for two neutrino flavours:

$$\Gamma \propto \frac{(\gamma_S + \gamma_D)/2\pi}{(E_{S,0} - E_{D,0})^2 + \frac{(\gamma_S + \gamma_D)^2}{4}} \cdot \left\{ 1 - 2s^2c^2 \left[1 - \frac{1}{2} (e^{-L/L_S^{\text{coh}}} + e^{-L/L_D^{\text{coh}}}) \cos \left(\pi \frac{L}{L_{\text{osc}}} \right) \right] \right\}$$

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In realistic cases: $L_{S,D}^{\text{coh}} \gg L^{\text{osc}} \Rightarrow$ Decoherence is not an issue.

Outline

- 1 The Mössbauer neutrino experiment
- 2 Oscillations of Mössbauer neutrinos: Qualitative arguments
- 3 Mössbauer neutrinos in QFT
- 4 Conclusions**

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Thank you!